

NASA CR 65996

N68-18824  
*alpha*

ANALYSIS OF POLAR CAP ABSORPTION EVENTS  
III. TIME RELATION OF RF-H $\alpha$  MAXIMUM INTENSITY  
FOR ALL CM BURSTS  $\geq 500 \times 10^{-22} \text{ W}(\text{m}^2\text{c/s})^{-1}$

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Report No. 00.865

12 October 1966

LTV ASTRONAUTICS DIVISION  
LTV AEROSPACE CORPORATION

Second Quarterly Report  
Prepared Under Contract NAS 9-4911  
Supplemental Agreement 1 with  
National Aeronautics & Space Administration  
Manned Spacecraft Center

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1.0      OBJECTIVES, RESULTS, AND CONCLUSIONS

During the analysis of flares of importance  $\geq 3$  it was found that if the time of the RF maximum flux at one or more of the frequencies 1420, 1500, 2000, 2800, 2980, 3000, or 3750 Mc/s preceded the time of the H  $\alpha$  flare maximum, there was a probability of 90% that the flare would not be followed by a PCA event. Conversely, the time of RF peak flux followed the time of the H  $\alpha$  flare maximum for 76.5% of the flares that were followed by a PCA event.

It was also found that:

87% of the flares that were followed by a PCA event the RF peak flux exceeded  $500 \times 10^{-22} \text{ W}(\text{m}^2\text{c/s})^{-1}$

while

75% of non-PCA flares the associated RF peak flux was less than  $500 \times 10^{-22} \text{ W}(\text{m}^2\text{c/s})^{-1}$

In order to increase the statistical significance, the study has been extended to all known cases where the peak flux at one or more of the selected frequencies in the decimeter and centimeter range from 1420 through 3750 Mc/s exceeded  $500 \times 10^{-22} \text{ W}(\text{m}^2\text{c/s})^{-1}$  associated with flares of importance  $\leq 2+$ .

This study added 180 new events to the 142 investigated in the previous study.

1.1      TIME RELATION OF FLARES, RF EMISSIONS AND KNOWN PCA EVENTS

We were able to find sufficient reliable data, including the start and time of maximum of the flare, the start and time of RF peak flux, and peak flux to determine with reasonable confidence the time between the peak RF flux

and the H  $\alpha$  flare maximum for 113 cases. We were able to determine with reasonable confidence that 97 of these cases were not followed by a known PCA event. Nine were followed by an important PCA event and seven by a small or questionable PCA. From this analysis we find that 78% of all flares with RF peak flux\*  $\geq 500 \times 10^{-22} \text{ W}(\text{m}^2\text{c/s})^{-1}$  that were not followed by a PCA, the time of RF peak intensity preceded or coincided with the time of the flare maximum.

In the case of the minor flares that were followed by a PCA event we find only 44% where the time of the RF peak flux follow the time of H  $\alpha$  flare maximum.

This is in strong contrast with flares of importance  $\geq 3$  where (ref. 13) it was found that 76.5% of the flares that were followed by a PCA event, the time of RF peak flux followed the time of the H  $\alpha$  flare maximum.

## 1.2 ANALYSIS OF FAILURES AND FALSE ALARMS

### 1.2.1 RF-H $\alpha$ Failures

An RF-H  $\alpha$  event is considered to be a failure if the event (with reasonable confidence) was followed by a PCA and the time of the RF peak intensity preceded the time of the H  $\alpha$  flare maximum.

### 1.2.2 RF-H $\alpha$ False Alarm

The RF-H  $\alpha$  event is considered to be a false alarm if the event is not followed by a known PCA and the time of RF peak intensity followed the time of H  $\alpha$  flare maximum.

The 21 false alarms from the present study and the six from the previous study (ref. 13) and the failures from the two studies will be examined

\* In our subsequent discussion we will use the term "flux units" where one flux unit is  $10^{-22}\text{W}(\text{m}^2\text{c/s})^{-1}$ .

in greater detail during the next quarter in a search for additional necessary and/or sufficient conditions.

2.0

RF EMISSIONS WITH PEAK FLUX  $\geq$  500 FLUX UNITS ASSOCIATED WITH  
FLARES WITH IMPORTANCE  $\leq$  2+

Table 1 lists all RF emissions reported by at least one radio observatory in the frequencies range 1420 through 3750 Mc/s, with a peak flux  $\geq$  500 flux units associated with flares with importance  $\leq$  2+.

A total of 180 events were compiled from all available sources (references 4, 9, 10, 14 and 23). After a detailed analysis of all available data the RF events were classified as follows.

	No. Events	Sub Total	Grand Total
No flare patrol	15		
No flare reported, flare patrol probable	11		
Time of flare maximum not reported	17		
Time of RF peak not reported	13		
Flare-RF association or RF peak questionable	11		
Total insufficient data	67		
Flares not followed by a PCA	97		
Flares followed by a PCA event	16		
Total events for detailed study	113		
Total events	180		

2.1

FLARES WITH MAJOR RF EMISSIONS NOT FOLLOWED BY A PCA EVENT

2.1.1

Delay Time  $\Delta t$  Between H $\alpha$  Flare Maximum and RF Peak Intensity

The time of all RF peak emissions reported by one or more of the seven frequencies tabulated in Table 1 were compared with the time of the H $\alpha$  flare maximum and a value for  $\Delta t$  was obtained (where  $\Delta t$  equaled the time

of RF peak emission minus the time of the H $\alpha$  flare maximum). The algebraic value for  $\Delta t$  was recorded for the emissions with peak flux  $\geq 500$  flux units only. In cases where a value could be obtained for more than one frequency, the value chosen for this analysis was restricted to only one frequency with preference given to 2800, 3750, or 3000 respectively. The number of events for each frequency that were chosen in this way, with the algebraic sign for  $\Delta t$  are shown in Table 1.1.

Frequency	$\Delta t < 0$	$\Delta t = 0$	$\Delta t > 0$	Total
1420	1	0	0	1
1500	3	0	2	5
2000	1	0	2	3
2800	20	2	8	30
2980	6	0	0	6
3000 Tok	16	3	5	24
3000 HHI	2	0	2	4
3750	17	5	2	24
	66	10	21	97

TABLE 1.1

Number of RF-H $\alpha$  Events Not Followed by a PCA.  
Analyzed at Each Frequency

The frequency distribution for RF-H $\alpha$  events followed by an important or a small PCA is shown in Table 1.2

Frequency	Important PCA		Small or Questionable PCA		Total
	$\Delta t \geq 0$	$\Delta t < 0$	$\Delta t \geq 0$	$\Delta t < 0$	
1420	1	0	2	0	3
1500	0	0	0	0	0
2000	0	0	0	0	0
2800	1	3	0	0	4
2980	0	0	0	0	0
3000 Tok	0	1	1	1	3
3000 HHI	0	0	0	0	0
3750	1	2	3	0	6
Total	3	6	6	1	16

TABLE 1.2

Number of Major RF-H $\alpha$  Events Followed by a Known or Questionable PCA Analyzed at Each Frequency

The results of the present study which extends the analysis to all minor flares (importance  $\leq 2+$ ) with a peak flux  $\geq 500$  flux units at one or more of the chosen frequencies are shown in Figure 1a and Table 1.1. We find that for 78% of all of the flares that were not followed by a known PCA event the time of the peak RF intensity coincided with or preceded the time of the H $\alpha$  flare maximum ( $\Delta t \leq 0$ ).

#### 2.1.2 Duration of the RF Bursts

The energy released at the time of a flare is directly proportional to the time integral of the RF flux.

The time durations together with the range of peak flux and the delay time between the RF peak intensity and the H $\alpha$  flare maximum are shown on Figure 2 for the three classes of flares.

- (a) Flares not followed by a PCA event.
- (b) Flares followed by an important PCA.
- (c) Flares followed by a small or questionable PCA event.

The data shown on this figure introduces a number of problems that will be discussed in detail in the next quarterly report.

These include:

- (a) The 21 flares that were not followed by a known PCA event where the time of the RF peak intensity followed the time of the H $\alpha$  flare maximum (the false alarms). In particular those events where the duration of the RF emission was greater than 20 minutes or the RF peak intensity exceeded 1000 flux units.
- (b) Flares that were followed by a PCA event, where the time of peak RF flux preceded the time of the H $\alpha$  flare maximum (the failures).

It was found (ref. 17) that all RF bursts at 2800 and 3750 Mc/s that were followed by PCA events for which time histories were available, the time integrated RF flux gave energies in excess of  $10^{-17}$  joules ( $m^2 c/s$ ) $^{-1}$ . Integrated RF energies were calculated for seven of the nine events (ref. 18) as shown in Table 2.1 below, with other data from Tables 1 and 2. We have also included the data for the event on 22 June 1957, for which a peak of 570 flux units was reported at 3000 Mc/s by Tokyo, a flux time history is not available at this time but an approximate energy value (peak flux x duration/2) is given. Bailey (ref. 2) gives an equivalent 30 Mc/s absorption of 5 db,

but indicated that this event might be a continuation of the weak event on the 19th. Neither Kahle (ref. 15) or Reid and Leinbach (ref. 21) give an absorption for the event on the 20th. It is quite probable that this was a small PCA event.

It is interesting to note that seven of the nine events occurred on the Eastern quadrant of the sun.

All of the events will be examined in detail in the next quarterly report.

Date	PCA Onset	Int.	RF Onset	Units Int.	Peak Dur.	$\Delta t$	Flare Max.	Imp.	CMD	Integrated Energy
6/19/57	2215	W	1609	2325	10	- 3	1613	2	E45	34
8/28/57	2400	3.2	2018	760	5	- 4	2024	2+	E30	10
8/20/58	21/1400	3.0	0042	1450	5	- 2	0044	2+	E18	15
9/03/60	0500	2.7	0039	12000	85	- 3	0108	2+	E88	160
9/10/61	2100	2.9	1930	880	61	- 9	2010	1	W80	95
6/22/57	0500	5.0	0231	570	21	- 3	0241	2	E12	(36)
2/09/58	10/0600	3.2	2138	856	32	+10	2142	2+	W14	
3/29/60	0800	2.6	0655	8250	52	+24	0710	2+	E30	915
3/30/60	2000	5.0	1518	640	22	--				
			1540	1750	60	+16	1540	2	E13	160
			1640	138						

TABLE 2.1

Summary of PCA Events Associated with Flares Importance  $\leq 2+$   
with RF Peak  $\geq 500$  Flux Units with Integrated Energy  
in Units  $10^{-18}$  Joules ( $m^2 c/s$ ) $^{-1}$

2.1.3    Central Meridian Distance of Flares Associated with Major RF Bursts

Figure 3 does not indicate any real preferential central meridian distance for flares associated with the major RF bursts. The seven events at E60 and E70, Figure 3b, will be examined in more detail with the five events in the eastern quadrant of the sun shown on Figure 3e.

2.1.4    Time Distribution of All Bursts  $\geq$  500 Flux

Figure 4 shows a reasonable correlation between the number of RF events reported and the number of radio observatories normally observing during each hour of the universal day. The number of events during the times of no flare patrol or no flare reported shows a strong concentration during the first three hours of the universal day. This is in agreement with the analysis by Dodson and Hedeman (ref. 6) who found the least number of flares of all importance during the first five and last five hours of the universal day.

Figure 4 also shows that there were several periods during the universal day when RF patrol coverage was marginal.

An important PCA is indicated by the letter P in the appropriate box. The S indicates a small or questionable PCA. The PCA at 0400 on April 5, 1960, is not included, since the association of the RF emission starting at 0140 with the importance 2 flare which started sometime before 0215 is considered to be questionable. It is probable that the H  $\alpha$  flare maximum occurred sometime before 0215.

## 2.2 FLARES WITH MAJOR RF EMISSIONS FOLLOWED BY A PCA EVENT

During the analysis of the RF-H  $\alpha$  time association for flares of importance  $\geq 3$ , it was found (ref. 13) that the time of the RF peak flux followed the time of H  $\alpha$  maximum for 76.5% of the PCA flares.

The use of this criteria for classifying an RF-H  $\alpha$  event as a probable PCA or a non PCA event is almost a complete failure (Fig. 1b) since only three of the nine important PCA are associated with a positive  $\Delta t$ . While in the case of the smaller questionable PCA events that could be associated with an RF-H  $\alpha$  event six of the seven (Fig. 1c) were associated with a  $\Delta t \geq 0$ .

Figure 1d shows the RF-H  $\alpha$  values for  $\Delta t$  for minor flares and RF emissions less than 500 flux units, where the values of  $\Delta t$  were positive for only three of the seven events.

The large percentage of failures associated with flares of importance  $\leq 2+$  may be due to questionable flare association, or failures to observe the true time of the flare maximum. This situation will be examined during subsequent investigations.

3.0 GENERAL DISCUSSION AND PROGRAM FOR THE NEXT QUARTER

3.1 COMPARISON OF PEAK FLUX REPORTED AT DIFFERENT FREQUENCIES

During the progress of this study it was noticed that a wide range of peak fluxes were reported at the different frequencies for a given event.

In particular it was found that for 47 of the events where both Tokyo (3000 Mc/s) and Nagoya (3750) reported peak fluxes, the peak reported at 3000 Mc/s was greater than the peak reported by Nagoya for 37 cases. In most cases the Tokyo values were greater than the Nagoya by a factor of at least two and in some cases by a factor of 10 or greater.

Figure 6 shows a scatter plot of the peak flux value at 3000 and 3750 Mc/s with the least squares fit. The reported times of the start of the emissions, and peak flux agree plus or minus a minute or so in all cases.

A corresponding comparison of peak fluxes at 2800 and 3750 Mc/s show an almost perfect correlation. Unfortunately, the total number of points is small since they were made near sunrise at 3750 and sunset at 2800. Only one event shows a large discrepancy (9/28/61 at 2211 UT) when a peak of 800 flux units was reported at 2800 Mc/s while a peak of 1690 was reported at 3750; Ottawa reported interference, consequently the real value of the peak flux may be much greater than 800.

Figure 7 shows a comparison of the daily mean flux for the first six months of the IGY. This shows an almost complete agreement between 2800 Mc/s and 3750 Mc/s while the Tokyo mean values average more than 100 flux units higher. However, the daily fluctuations at all of the frequencies are in complete agreement.

Figure 8 shows a similar comparison of the daily mean flux for last six months period of 1960 during the decline of solar activity from the

maximum toward minimum. It is seen that the close correlation of the flux at 2800 and 3750 Mc/s still exists while 3000 Mc/s remains high.

### 3.2 PROGRAM FOR THE NEXT QUARTER

#### 3.2.1 Association of H $\alpha$ and Other Parameters with RF Time Histories

Wherever possible the time histories of the H  $\alpha$  -RF events will be examined in greater detail and the possibility that other minor flares preceding or during the time of the event may influence the classification of the event as a failure or false alarm or a questionable association will be investigated.

#### 3.2.2 Investigation of Sunspot Characteristics

Several investigators have carried out a limited study of sunspot magnetic field configurations and have found a high probability that a proton producing flare will occur in sunspots with umbra of opposite polarity in the same penumbra (classified as a  $\delta$  magnetic configuration). Sunspot groups associated with both RF-H  $\alpha$  failures and false alarms will be examined in all cases where Mt. Wilson sunspot drawings and magnetic field measurements were made.

It is anticipated that the many gaps in the Mt. Wilson data will greatly restrict the number of events that can be evaluated, and an effort will be made to fill the gaps with data from other observatories.

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TABLE I

RADIO EMISSIONS AT CENTIMETER WAVELENGTHS WITH PEAK  
 $\text{FLUX} \geq 500 \times 10^{-22} \text{ W}(\text{m}^2/\text{s})^{-1}$  ASSOCIATED WITH FLARES OF IMPORTANCE  $< 3$  DURING THE 19TH SOLAR CYCLE

DATE	ASSOCIATED SOLAR FLARE					SHORT WAVE FADE			SPECTRAL				SINGLE FREQUENCY 1500 TO 3750 Mc/s								PCA No.		
	Start Min.	Duration	Time	Position	Imp.	On-set Min.	Dur.	Type	Imp.	Start/Dur	Imp.	II	IV	Time	Duration	Peak	*** Int.	At	No.	Int.			
1955																							
1.	2/24	No flare patrol	0000 to 0500			None reported			0104					#3750	Mag	CD	0101.5	0102.5	2.5	1470	P		
2.	6/18	1904	36	1910	S22 W21	2+(3+)		1903	47	S	2+			#2800	Ott	SD	1906.5	1907.8	30.5	1580	-2.2		
3.	11/15	No flare patrol	1600 to 1800 UT			None reported								#2800	Ott	CA	1734	1737	5.0	565	P		
1956																							
4.	2/16	1805	154	--	N22 E06	2+		1802	93	SL	3			#2800	Ott	CD	1756	1813	51.0	623	M <sub>1</sub>		
5.	2/19	1430	147	1445	N25 W23	1+(2)		1429	151	S	3			#2800	Ott	SD	1425	1435	29.0	643	-10		
6.	3/01	1730	15	--	N20 W55	2		1726	22	S	2			#2800	Ott	SD	1721	1724	5.0	614	M <sub>1</sub>		
7.*#3/10	0515	85	--	H16 E88	2		0438	117	SL	3-			#3000	Tok	CD	0443	0518	80.0	850	M <sub>1</sub>	2 3.5		
8.	3/13	1453	32	--	N21 E50	2		1452	118	S	3-			#2800	Ott	SD	1451	1454	14.5	860	M <sub>1</sub>		
9.	3/15	1617	80	1635	N22 E21	2+(3)		1623	120	S	3			#2800	Ott	SD	1621	1627	23.5	1320	-8		
10.	5/18	1600	10	--	S26 E45	1-		1605	15	S	1			#1500	HII								
														#2800	Ott	SD	1604.3	1604.4	15.7	515	M <sub>1</sub>		
11.	5/30	No flare reported Probably no patrol						0230	95	S	3+			#3000	Tok	CD	0232.5	0235	14.5	600	P		
														3750	Hag	CD	0234		7.3	440			
12.	7/22	1624	56	1641	N29 W54	2 (2+)		1635	110	S	2+		(1638)	#2800	Ott	SD	1638	1641.8	8.5	660	+0.8		
13.	11/08	0231	17	0242	N14 E62	1		0243	17	S	1			3000	Tok	CD	0223	0224	8.0	369			
														3750	Hag	CD	0238	0245	10.0	538	+3		
14.	11/22	1312	63	1341	S15 W83	2		1330	65	S	3		(1323)	#2800	Ott	SA	1333	1334.7	3.0	17			
														*	SA	1336	1339	20.3	1000	-2.0			
15.	12/20	0432	15	0459	N12 E15	1+(2)		0442	81	S	3		(0444)	#3000	Tok	CD	0444	0451	24.0	530	-8		
16.	12/26	1401	67	1412	S17 W11	2 (3)		1403	97	SL	3-		(1403)	#2800	Ott	CD	1403	1454	165.0	800	+42		
17.	12/29	0040	100	0045	N16 E59	1+		0044	106	S	3+		(0043)	#3000	Tok	CD	0043	0056	90.0	1150	+11		
1957																							
18.	1/05	No flare patrol from 1500 UT on the 4th to 0125 on the 5th						0050	63	SL	2+		(0050)	#3000	Tok	CD	0050	0056	58.0	501	P		
19.	1/06	No flare patrol between 1500 and 2100 UT						1702	53	S	2+		1703 9/3+	1711 169/3	#2800	Ott	SD	1701.5	1703.5	10.0	700	P	
20.	1/24	1638	15	--	S28 W80	2		1638	27	S	2+			#2800	Ott	SD	1637	1658.8	10.0	1000	M <sub>1</sub>		
21.	2/08	1550	25	1555	S28 E38	2		1552	10	S	2		1551 4/3+	#2800	Ott	SD	1550	1551	6.0	865	-4		
22.	4/02	0255	109	0339	S16 W44	2		0250	120	G	3			#3000	Tok	CD	0301	0336.5	60.	800	-2.5		
23.	4/12	1850	80	1920	S25 W73	2(2+)		1856	89	S	3+		1904.7 11.3/3	#2800	Ott	SD	1855.5	1900.5	23	525	-19.5		
24.	4/18	1310	43	1323	S16 E14	2		1304	36	S	2+		1304 8/3	(1305)	1500	HII	CD	1305	1305	25	451		
														2800	Ott	SD	1304		6.0	550	M <sub>2</sub>		
25.	6/04	0859	41	0902	S17 W27	2		0900	30	S	3-		(0859)	*2980	Ned	SD	0859	0917	10	350			
														*	SD	0928		11	610		M <sub>2</sub>		
26.	6/05	1326	67	1330	S17 W3	2		1328	26	S	3-		1329 4/3+	#2800	Ott	SD	1327	1328.2	8.0	725	-1.8		
														#2980	Ned	SD	1325		4.0	>670			
27.	6/06	1130	18	1133	S14 W27	1		None reported						#2800	Ott	SD	1129	1129.8	1.5	525	-3.2		
														#2980	Ott	Ned	CD	1129		1.0	1700		
28**	6/19	1609	40	1613	N20 E45	2(2+)		1608	44	S	3		1615 5/3	(1609)	#2800	Ott	SD	1609	1610	10.0	2325	-3.0	7 W
														#2980	Ned	SD	1625		4.0	>670			
29**	6/22	0236	21	0241	N23 E12	2		0229	74	S	2		(0231)	#3000	Tok	CD	0233	0238.4	8.0	90	-2.6	8 5.0	
														#3000	Hag	CD	0231	0238.4	21.3	570	-2.6		
30.	6/27	2322	56	2335	N20 W62	1		2325	55	SL	1		(2408)	1420	Syd	CD	2357		8.2	>239			
														#3000	Tok	CD	2408		25.0	504			
														3750	Hag	CD	2411		16.0	42			
31.	7/02	No flare reported						0013	47	SL	2		(0015)	2000	Hag	ECD	0016	0016.3	2.5	70			
														2800	Ott	SD	0015	0017	6.0	180			
														#3000	Tok	CD	0015	0016.5	23.0	>330			
														3750	Hag	SD	0015	0016.3	3.0	305			
32.	7/21	0633	87	0658	N30 E15	2		0647	60	S	3			#2980	Hag	SD	0659.5	--	7.0	536			
														#3750	Hag	ECD	0633	0633.5	1.5	106			
33.	7/21	1320	82	1337	N29 E12	2(3)		1335	45	S	2+		(1321)	2800	Ott	CA	1321.5	1322.5	6.5	23			
														*	SD	1334.5	1335.9	8.0	850	-1.1			
34.	8/10	0125	17	0129	N26 W71	1		0100	60	SL	3		0129.4 44/3	1420	Syd	ECD	0128	--	2.5	>196			
														#2000	Hag	SD	0127	0128.6	3.0	550			
														#3000	Tok	CD	0126	0127.5	2.5	1400			
														#3750	Hag	ECD	0126	0127.4	4.0	1700	-3.0		

#### **REFERENCES**

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## THEORY OF STATE FUNCTIONS

WITH PCR ASSOCIATION

$\mu_2$  times of  $\lambda\gamma$  peak unknown

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TABLE I (Continued)

DATE	ASSOCIATED SOLAR FLARE							EFFECTIVE			SPECTRUM (WAVELENGTHS)						PCA						
	Start	Duration	Time	Position	Imp.	On-set	Dur.	Type	Imp.	Start/Dur/Imp.	II	IV	f	Obs.	Type	Onset	Time	Duration	Peak Int.	At	No.	Int.	
92. 6/26	0245	152	0306	N10 W49	2+ (3)	0247	85	SL	2+	0306 (0255) 16/-	2000	Mag	CD	0244	0245.9	4.0	34						
							0319	Tok	SD	0243	2000	Mag	CD	0245	0245	7.0	410						
							110/1	Mag	CD	0244	3750	Mag	CD	0244	0244.9	3.0	57						
										1420	Syd	SD	0255	0316	30.0	189							
										2000	Mag	SD	0255	0316.2	30.0	100							
										*3000	Tok	SD	0255	0307.8	22.0	507	R <sub>6</sub>						
										3750	Mag	CD	0255	0307.8	30.0	155							
93. 6/27	0254	71	0308	N10 E37	1+ (2)	0308	25	S	2		2000	Mag	SD	0306	0306.8	3.0	70						
										*3000	Tok	SD	0306	0306.5	6.0	593	- 1.5						
										3750	Mag	SD	0305	0306.7	7.0	245							
94. 7/03	0041	33	0050	N30 E37	1+ (2)	0050	7	S	1-		1420	Syd	SD	0048	0048	2.0	206						
										2000	Mag	SD	0046	0048.2	3.0	125							
										*3000	Tok	SD	0045	0048.2	6.0	620	- 1.8						
										3750	Mag	SD	< 0048.1	0048.1	>1.0	310							
95. 7/04	0513	21	0517	N29 E26	1+	0517	11	S	1		1420	Syd											
										2000	Mag												
										*2980	Med												
										*3000	Tok	CD	0516	0517	4.0	>789					+0.1		
										*3750	Mag	SD	0515	0517.1	2.5	800							
96. 7/20	No flare reported							0244	31	S	1		2000	Mag	SD	0241	0242.6	2.5	115				
										*3000	Tok	SD	0241.0	0242.5	4.0	597	P						
										3750	Mag	SD	0241	0242.5	2.5	230							
97. 7/24	2327	121	2443	N10 E85	2+	2444	56	G	2		2000	Mag	SD	2440	2443.2	6.0	(75)						
										*3000	Tok	SD	2441.3	2443	7.0	828							
										*3750	Mag	SD	2441	2443.1	4.0	535	+0.1						
98. 7/29	0458	28	0458	S14 W38	1	None reported							2000	Mag	CD	0503	0510	40.0	(31)				
										*3000	Tok	CD	0505.2	0512	8.0	506	+14						
										3750	Mag	SD	0503	0510	30.0	(10)							
99. 8/02	1840	11	1841	S14 W90	1-	1840	153	S	3+	1843 (1840) 6/3	2800	Ott	SD	1840	--	>300.0	30						
										2980	Med	SA	1840	1842.1	25.0	2050	+1.1						
100. 8/20	0042	46	0044	N16 E18	2+ (3)	0042	33	S	2+	0046 (0042) 5/3	1420	Syd	CD	0040	0044	13.0	255						
										*2000	Mag	ED	0040	0044	7.0	(620)	R <sub>7</sub>						
										*3000	Tok	CD	0041	0043.3	7.0	1260							
										*3750	Mag	ED	0041.5	0042.4	5.0	1450	-1.6						
101. 8/28	1025	20	1030	S18 W64	2+	1023	62	S	3	(1028)	1500	HII	CD	1009	1039	53.5	215						
										2800	Ott	CD	1028	1039.5	20.0	350							
										* 3000	HII	CD	1019	1039	36.0	573	+9.0						
										HII also max. at 1028 and 1030 OTT in sunrise oscillation													
102. 9/14	0822	128	0835	S10 W80	2+ (3+)	0851	58	S	3	(0832)	1500	HII	CD	0834	0904	85.0	460						
										*2980	Med	SD	0833.5	--	5.0	371	R <sub>9</sub>						
										*3000	Tok	CD	0847.5	--	43.0	870							
										*3750	HII	M	0830	0904	69.0	1259	+29						
103. 10/19	0634	106	0725	S17 W35	2+ (3)	0720	30	S	2		1420	Syd											
										2000	Mag												
										*2980	Med												
										*3750	Mag	ED	0723	0723.9	2.0	585	-1.1						
104. 10/21	2318	129	2330	S04 W22	2+ (3)	2328	72	S	3+	2328 (2327) 13/3+	2000	Mag	CD	2323	2327	55.0	428						
										*2328	2344	SD	2355	2345	520								
										*3750	Mag	CD	2323	2327	55.0	1150	-3.0						
105. 11/21	No flare patrol 2230-2430							None reported							*1420	Syd	2318	2320	3.0	532			
										2000	Mag	CD	2316	2319.1	5.0	(165)	P						
										*2980	Med	CD	2320.4	2320.4	3.0	(25)							
106. 12/11	1116	37	1127	S01 E03	2	1122	22	S	2	(1805)	*2800	Ott	SD	1805	1810	20.0	1225	-2.0					
										*2800	SD	SD	1805	1810	20.0	1225							
107. 12/12	1802	40	1812	S02 W06	2	1808	32	S	2+	(1805)	1500	HII	SD	1251.5	1252.8	2.5	191						
										*2800	SD	SD	1252.8	1252.8	1.0	35							
										*2980	SD	SD	1257	1300.5	18.0	1500	-3.5						
										*2980	Med	CD	1258	--	18.0	~1799							
109. 12/17	1047	28	1053	S16 W57	1	None reported							1500	HII	SD	1039	1041	19.0	420	-12.0			
										*2980	Med	SD	1039.5	--	5.0	571							
										*1420	Syd	F	0559	0605	11.0	165							
										*2000	Mag	CD	0534	0605	65.0	370							
										*3000	Tok	CD	0536	0605	>50.0	1750	-19.0						
										*3750	Mag	CD	0534	0605	50.0	1020							
110. 12/23	0543	138	0624	S15 W66	2+ (3)	0540	73	G	3+	(0559)	1500	HII	SD	0946	0947	6.5	260						
										*2980	SD	SD	0946	--	3.0	493							
										*3000	HII	SD	0946	0947	10.5	470	-3.0						
										Jod	ED	0950	0951	50.0	330								
111. 12/24	0946	24	0950	S14 W54	1 (1+)	0943	17	S	1		1500	HII	SD	0946	0947	6.5	260						
										*2980	SD	SD	0946	--	3.0	493							
										*3000	HII	SD	0946	0947	10.5	470	-3.0						
										Jod	ED	0950	0951	50.0	330								
	<u>1959</u>									*2800	Ott	--	2130	--	15.0	2000	R <sub>2</sub>						
112. 1/14	2140	12	2142	N21 K11	1-	None reported							*1500	HII	CD	1342	1342.5	7.0	645	+22			
113. 2/08	1342	94	1410	N26 W69	1 (1+)	1342	29	S	2	1344.5 (1342) 5.3/3	2800	Ott	CD	1342.5	1343	8.0	(180)						
										2980	Med	CD	1340.5	1									

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	DATE	ASSOCIATED SOLAR FLARE				SHORT WAVE FADE			SPECTRAL		SIMPLE FREQUENCY 1500 TO 3750 Mc/s						PCA					
		Start Min.	Duration	Time	Position	Imp.	On-set	Dur.	Type	Imp.	Start/Dur/Imp.	II	IV	f	Obs.	Type	Onset	Time Max.	Duration Min.	Peak Int.	Alt	No. Int.
148**	3/30	1455	339	1540	M12 E13	2 (3+)	1520	160	SL	3	1529 11/3	1526 454/3	*2800	Ott	C+	1518	1527 1556	22.0 196.0	640 1750	+16		40 5.0
149.	4/03	< 0317	5	0317	M12 W33	2	0305	25	S	2+			*2000	Mag	C	0306	0310.5	5.0	400		- 6.4	
150.	4/05	< 0215			M12 W62		0140	157	SL	3+	0152 15/3	0207 53/3	*2000	HFI	C+	0140	0206.1	125.0	1230			42 3.1
151.	4/29	No flare reported				None reported							*2000	Mag	C	0525	0538.1	23.0	990			
152.	5/26	0850	120	0928	M14 W15	2+	0914	46	S	2		(0909)	*2800	Ned	C+	0909	--	27.0	1350	M2		
153.	6/25	< 1026	> 20	1029	M19 E03	1+ (2)	1027	33	S	2			1500	HFI	S	1025.8	1027.1	10.5	193			
154.	6/25	2039	61	2046	M19 W04	2+ (3)	2040	30	S	2-	2048 17/3	2045 66/3	*2800	Ott	C+	2037	2046	40.0	700	0		
155.	6/29	0125	82	0148	M21 W50	1 (1+)	0138	128	S	2			2000	Mag	F	0137	0142.3	17.0	240			
156.	6/30	1029	76	--	M18 W64	1 (1+)	1030	92	S	2-			*1500	HFI	C	1019.5	1037.1	52.5	511			M1
157.	8/07	0724	41	0737	M18 E84	1	0730	55	S	2+			2000	HFI	C	0723.5	0730.2	50.5	288			
158.	8/11	0223	97	0257	M21 E35	2 (2+)	0225	90	SL	2			2000	Syd	F	0223	0235.0	35.0	375			
159.	8/11	1916	99	1929	M22 E26	2+ (3+)	1925	65	S	2	1920 9/3+	1926 53/2	*2800	Ott	S	0916	0928	144.0	9			
160.	8/14	0511	104	0525	M22 W06	2+ (3)	0515	45	S	3			*2000	Mag	C	0515	0518.2	20.0	775			
161.	8/14	1242	92	1310	M20 E36	2 (2+)	1307	53	S	3-			2800	Ott	S	1307	1310.7	16.0	680	+ 0.7		
162**	9/03	0037	77	0108	M18 E88	2+ (3)	0045	126	SL	3+	0102 22/1	> 0038 16/2	*2000	Mag	C	0035	0105.2	90.0	7100			48 2.7
163.	9/16	1710	105	1724	S22 E68	1	1709	101	S	3	1714 14/3	1717 114/3	*2800	Ott	C+	1702	1712	23.0	550			M11
164**	9/26	0525	51	0539	S22 W64	1+ (2+)	0520	121	SL	3+	0543 21/2	< 0554 17/1	1420	Syd	C	0529	--	38.0	> 139			
165.	10/10	0710	86	0722	S17 E23	1+ (2)	None reported						2000	Syd	C	0704	0718	21.0	30			
166.	10/11	0517	159	0535 0600	S17 W36	2	0525	63	S	3	0530 17/3	> 0532 40/2	2000	Mag	C	0523	0527.6	26.0	630			
167**	11/11	0305	83	0340	M28 E12	2 (2+)	0311	185	S	3+	0349 8/2	0330 219/3	1420	Syd	C	0316	0320	120.0	46			G32 VS
168**	11/14	0246	124	0304	M27 W20	2+	0300	120	SL	3	0305 115/3		*2000	Syd	FCS	0258	0443.7	140.0	1800			G:34 VS
							(Gregory associates a proton event on 11/14 at 22xx UT with an importance 1+ at 2114 UT. The importance 2+ flare at 0246 seems to be a more logical candidate.)															
169.	9/10	1950	64	2010	M08 W80	1	1942	101	SL	3	1947 27/3	1937 50/3	*2800	Ott	C	1930	2001	61.0	880	- 9.0		56 2.9 N13
170.	9/16	1057	121	1110	M18 E77	2+ (3+)	1102	50	S	2			1500	HFI	C	1055.0	1104.0	70.0				
												*2900	Ned	C	1101.0	1111.6	34.0	532	- 5.4			
171.	3/13	1444	116	1446	M10 E66	2+	1448	94	S	3			*1500	HFI	C	1448.7	1451.6	70.0	> 880	+ 5.6		
												2000	HFI	C	1447.5	1450.3	19.5	470				
172.	1/30	1305	70	1306	M10 W54	1	None reported						1500	HFI	C	1322.5	1439.5	100.0	877			N14
173.	5/01	0525	190	0608	M15 E46	2	0530	39	SL	2+			2000	HFI	C	0526	0545	40.0	600			
174.	9/14	2112	53	2123	M12 E72	1	None reported						*3750	HFI	C	0526	0544	34.0	1500	- 24		
175.	9/14	2202	19	2206	M12 E73	1	None reported						2800	Ott	C	2230	2235.5	12.0	1880			
176.	9/15	0015	124	0042	M15 E75	2	0015	180	S	3+	0027 22/2		*2000	Mag	C	0017	0028	110.0	210			N15
177.	9/16	1430	62	1505	M12 E48	2	1440	125	G	3			3750	Mag	C	0015	0028	90.0	650			
178.	9/18	2236	144	--	M12 E20	1	None reported						2000	Mag	C	2231	2327.0	130.0	1100			M1
179.	9/20	2314	167	2403	M10 W09	2	2351 214	8	S	3	2400 76/2		2000	Mag	C	2350	2358.4	15.0	1200			
180.	10/28	0135	120	0158	M12 W24	3	0140	140	SL	3			*3750	Mag	C	0139	0156.3	40.0	580	- 1.7		
												2000	Mag	C	0156.3	0156.3	50.0	465				

**TABLE II**  
MUCH FLARES FOLLOWED BY A PCA EVENT

Date	FLARE UT Min.	UT Sec.	Position Flare	PEAK CIV ABSORPTION			HELIOT WAVE PACE			SPECTRAL			NOTE
				Start	Dur.	Abs.	Start	Dur.	Obsr.	Intens.	Energy	Peak	
<b>1956</b>													
1	3/10	0515	95	-- 2	H16 826 3431	2 3 <sup>b</sup> 0000	160	3.5	0430	117	3+		
2	11/13	1430	05	150L 2	H16 910 3753	4 5 <sup>b</sup> 2000	63	5.4	1430	120	2+		
									3000 CD	0441		80	890
									3750	0447	0518	80	1000
									2000 BD	1420		118	-41
									2000 BD	1431	1440	118	256
									2000 BD	1431	1440	118	175
									2000 BD	1431	1440	118	-21
													No emissions reported at meter wavelengths. Also 9400 Mc/s at 1433 UT flue (3P).
<b>1957</b>													
3	6/19	1608	941	1613 2(2+)	H20 845 4724	7 8 <sup>b</sup> 0000	2215	Week	1608	144	3	1615 5/3	--
4	6/22	0836	21	0841 2	H23 812 4029	8 2 <sup>b</sup> 2000	115	5.0	0829	78	2		
5	8/09	0617	63	0620 2	H09 876 4099	11 9 <sup>b</sup> 1600	50	3.1	0615	35	3-		
6	8/28	2010	38	2024 2+(3)	H20 830 4125	13 14 <sup>b</sup> 2000	27	3.2	2020	18	2+	2022 4/3	
7	8/29	1031	39	1036 2(1)	H25 820 4125	14 15 <sup>b</sup> 1300	58	8.2	1039	16	1+		
8	{ 09/08	1257	49	1303 1(2)	H10 826 4124	16 17 <sup>b</sup> 1700	46	7.2	1259	>25	0.1		
	{ 09/08	1313	317	1316 2+(3)	H34 836 4125	13 14 <sup>b</sup> 1700			1324	43	8 2-		
									2000 BD	1324-1	1324	8	+3
									2000 BD	1324-3	1324	8	+8
<b>1958</b>													
9	2/09	2108	114	2132 2+	H12 818 4400	22 2 <sup>b</sup> 0000	37	5.2	2124	20	SL 1		
10	3/25	0557	29	0603 2	H15 850 4476	25 2 <sup>b</sup> 1530	122	10.0	0603	27	SL 2	--	--
									3000 BD	0558.8	0559.3	>3.0	856
									3750 CD	0558.8	0559.3	10.0	100
													+10
11	6/06	0436	98	{ 0448 2(3)	H16 878 4578	25 1 <sup>b</sup> 0600			0436	50	2	0436 5/3	
				{ 0450									
									1420 CA	0445	0446	27.0	343
									2000 CD	0431	2135	29.5	207
									2000 CD	2135	2135	29.5	205
									2000 CD	0454.5	0455.7	4.0	33
									3750 BD	0505.5	0505.7	6.0	56
													+2.2
12	8/20	0042	66	0044 2+(3)	H13 817 4708	29 37 <sup>b</sup> 1800	19	3.0	0042	33	2+	0046 19/3	
									1420 BD	0040	0044	13.0	255
									2000 BD	0040	0044	7.0	580
									3000 BD	0041.2	0043.3	7.0	1260
									3750 BD	0041.2	0043.3	5.0	1450
													-0.7
													-1.6
13	9/22	0738	98	0750 2(2+)	H10 842 4705	32 4 <sup>b</sup> 1800	80	5.0					
									1420 BD	0737	0745	7.5	207
									2000 BD	0737	0745	6.0	356
									2000 BD	0739	--	31.0	356
									2000 BD	0733	0745	31.0	314
									3750 CD	0741	0746	20.0	45
													-5
<b>1959</b>													
14	3/29	0540	340	0710 2+(3)	H13 830 5015	39 1 <sup>b</sup> 0800	73	2.6	0552	121	3+		
15	3/30	1455	335	1533 2(3+)	H12 813 5015	40 0 <sup>b</sup> 0900	36	5.0	1520	180	3		
									1520	1526	1555	90.0	810
									1520	1518	1556	20.0	1750
									2000 CD	0655	0733.4	120.0	650
									2000 CD	0655	0733.5	52.0	650
									3750 CD	0655	0733.5	52.0	650
													+23.5
													+22.15
													+23.10
16	4/05	0415	105	0445 2(3)	H10 860 5115	40 1 <sup>b</sup> 0800	55	3.1	0440	187	3+	0452 15/3	
									2000 CD	0440	0456.5	120.0	1230
									3000 CD	0440	0456.7	64.0	2400
									3750 CD	0440	0456.7	64.0	6000
													M <sub>1</sub>
17	4/20	0107	441	0205 2+(3)	H14 821 5162	44 1 <sup>b</sup> 0900	134	34.0	0205	175	2+	0214 16/2	
									2000 C+	0205	0207.1	20.0	75
									2000 C+	0205	0207.1	20.0	185
									2000 C+	0205	0207.1	20.0	370
									2000 C+	0139	0206.0	9.0	115
									2000 C+	0139	0206.0	7.0	137
									2000 C+	0139	0206.0	7.0	37
									2000 C+	0139	0206.0	55.0	365
18	9/03	0037	77	0108 2+(3)	H18 888 5239	48 1 <sup>b</sup> 0500	89	2.7	0045	126	3+	0308 22/1	
									2000 C+	0305	0305.2	30.0	7100
									3000 C+	0305	0305.2	50.0	5800
									3750 C+	0305	0305.2	50.0	12000
													-2.6
													-3.4
19	11/20	*1955	37	2020 1(3)	H20 890 2124	51 1 <sup>b</sup> 0800	51	3.0	2023	82	3+	2028 17/2	
									2000 C	2023	2026.5	47.0	400
									2000 C	2023	2026.5	47.0	25
													+6.5
<b>1961</b>													
20	9/10	1555	1610 1	H10 890		2000 2100 2200	79	2.9	1942	101	3	1942 27/3	
									2000 C	1942	2021	61.0	880
									2000 C	1942	2021	61.0	95
													-9.1
21	9/20	2314	167	2403 2	H10 892 5264	16 1 <sup>b</sup> 4000	96	4.6	2351	214	3	2400 26/2	
									2000 C	2351	2356.4	15.0	1200
									2000 C	2351	2356.4	15.0	1200
									2000 C	2351	2356.4	15.0	1200
									2400 C	2351	2400.4	23.0	2100
									2400 C	2351	2400.4	23.0	2100
									2400 C	2351	2400.4	23.0	2100
													-4.6
													-27.5
													+7.5
													-2.7
													+29.7
													+72.7
<b>1962</b>													
22	9/20	2314	167	2403 2	H10 892 5264	16 1 <sup>b</sup> 6000							

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TABLE 3  
VERY SMALL PCA EVENTS ASSOCIATED WITH IMPORTANCE  
<3 FLARES AND RF EMISSIONS <500x10<sup>-22</sup> W(m<sup>2</sup>/c)-1

Date	Start	Duration	Max.	Imp.	Position	PCA No.	Start	Int.	SWF			II	IV	f	Type	Onset	Max.	Dur.	Int.	at			
									Onset	Dur.	Imp.												
1956																							
4/27	1546 2050	60	2100	1+	S14 E14 N16 W27	S2	2000	Weak	None reported	2053	24 SL	1+			2800	SD	1544 2051	--	2056	5.0	375		
1957																							
4/05	(1433 1444 1511 0550	13 3 15 135	-- 1 1 2	1 1 1 2	S15 W90 S20 E35 N14 E50 N23 W72	S4	06/0800	3.2	1408	32 SL	2				2800	CD	1407	1411.5	12.0	50			
4/06									None														
4/11	1722	88	1738	2+	S23 E04	S5	12/0120		1731	64 S	3				2800	SD	1725 CA	1729.5 1733	75.0 16.0	135	-5		
7/28	1346	72	1405	2	S23 W82	S7	1500	Weak	None reported						2800	SD	1353.2	1354.1	40.0	18	-10.9		
9/22	0558 0636 0643 0732	77 49 17 60	0614 0652 -- 0750	1+ 1 1+ 2	N10 W13 N08 W34 N24 W32 N23 W38	S8	1000								2000	CD	0643	0650	10.0	42	-2		
									0746	74 S	2+				2980	SD	0645	--	10.0	400			
															3750	CD	0642	0650	11.0	174	-2		
11/04	0558 0937 1058 1735 1949	18 27 26 15 10	0102 0200 -- 1740 1949	1 1 1 1- 1-	S20 W38 S12 W20 S24 W39 S25 W45 N23 W58	S9	05/0030 0200	2.6	None reported						2980	SD	1059.5	--	1.5	250			
1958																							
3/10	2024	64	2034	2	S12 W50	S12	11/0400	Weak	2025	30 0	1				2800	SD	2024 2028 3000 3750	2035 2032.5 0000 0021	90.0 7.0 30 38.0	10			
3/11	0030	>12	0034	1	N11 E02				0048	152 5	3					CD	0000	0000	180				
																CD	0021	0024.8	406				
3/14	0816 0920 1000 1136 *1454	54 21 60 24 47	0846 0927 1020 1143 1507	1+ 1+ 1 1 2	N07 W18 N08 W19 N07 W22 N07 W19 S21 W85	S13	2200	Weak							1500	CD	1453	1455	--	137			
									1455	130 SL	3				2800	CD	1458.5	1501	13.0	210	-6		
															3000	CD	1458	1501	34.0	375	-6		
1959																							
9/01	1359 1419 1422 1648 *1923 1947	51 31 53 104 173 43	1410 1424 1431 1704 1938 1953	2 1 1 2+ 2+ 1+	S12 W51 N10 W12 N14 W68 S12 W52 N12 W60 N09 W15	S18	02/0400	VS								SD	1418	1440	45.0	9			
																SA	1439.3	1420	2.0	30			
																SD	1645	--	146.4	20			
																CA	1658	1706	18.0	70			
																*BD	1928	2023	>212.0	50	+45		
																*CA	1932	2009	41.5	45	+31		
1960	None																						
1961																							
11/10	1121 1328 *1434	23 37 16	1136 1332 1444	1 1 1+	N09 W00 N09 W00 N19 W90	S21	1500	2 db							1433 36/3	1440 63/3	1500 2800	C C	1430 1428 1430	1435.3 1444 1433.5	10.0 38.0 10.0	~47 124 11	-8.7 0
1962																							
2/01	0901 1230 1522 1534	263 16 33 54	0907 -- 1556 1649	2 1 1 1	N10 W36 N10 W34 N09 W38 N09 W38	S22	2030	1-2 db							1657 36/3	23 S	1650 1636	S S	1550 1640	1554 20.0 5.0	20.0 38.0 10.0	3	8
1963																							
4/15	1034	116	1125	2	S11 W06	S23	1200	Moderate	1124	16	S	2	No observations		1500 2800	C S	1123 1123	1123 1124.5	1124.7 9.0	47.0 220	50 -0.5	-0.3	

\* Probable Flare

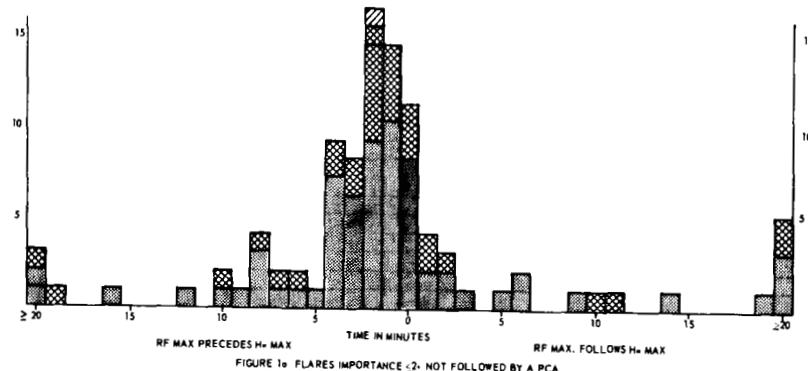


FIGURE 1a FLARES IMPORTANCE  $\leq 2$ , NOT FOLLOWED BY A PCA

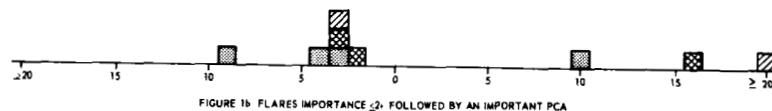


FIGURE 1b FLARES IMPORTANCE  $\leq 2$ , FOLLOWED BY AN IMPORTANT PCA



FIGURE 1c FLARES IMPORTANCE  $\leq 2$ , FOLLOWED BY A SMALL OR DOUBTFUL PCA

FIGURE 1 DELAY TIME ( $\tau$ ) BETWEEN H<sub>α</sub> FLARE MAXIMUM AND RF PEAK INTENSITY  
FLARES IMPORTANCE  $\leq 2$ , WITH RF PEAK FLUX  $\geq 500 \times 10^{-22} W (m^2 c/s)^{-1}$   
PERIOD 1954 THROUGH 1963, RF FREQUENCY 1420 THROUGH 3750 Mc/s.

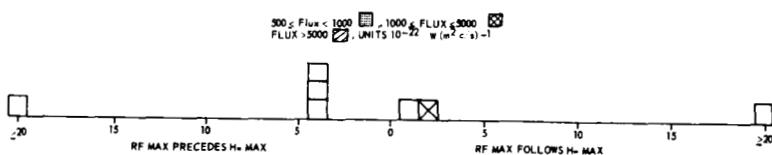


FIGURE 1d FLARES IMPORTANCE  $\leq 2$ , FOLLOWED BY A PCA, FLUX  $< 500 \times 10^{-22} W (m^2 c/s)^{-1}$

Flux < 250,  $\square$ ;  $250 \leq$  Flux < 500,  $\times$

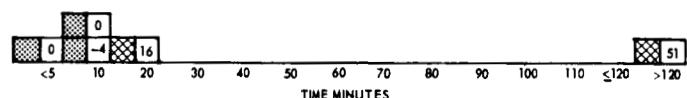
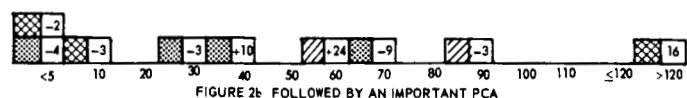
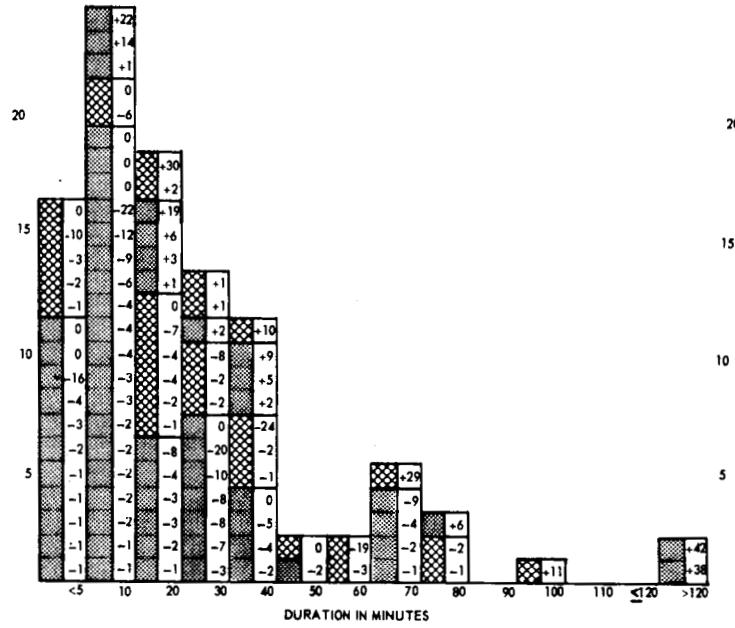
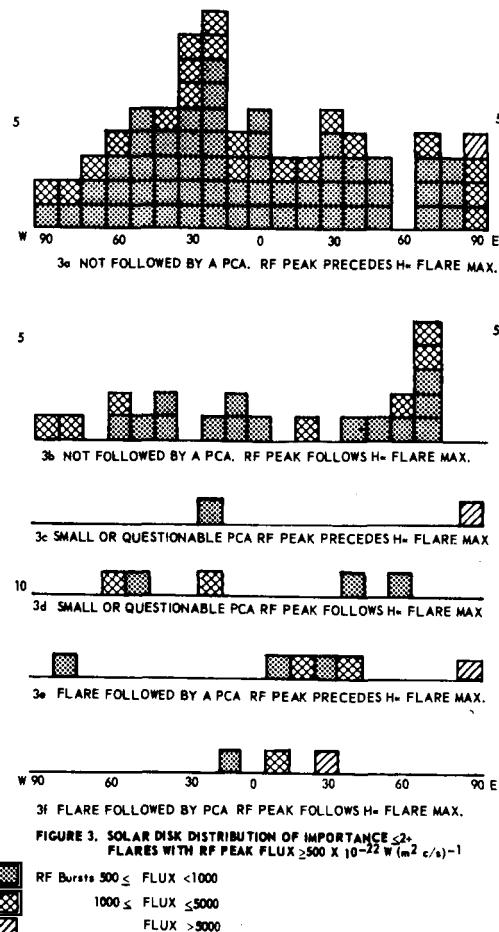


FIGURE 2 DURATION OF RF BURSTS FLARES IMPORTANCE  $\leq 2+$  WITH RF PEAK FLUX  $> 500 \times 10^{-22} \text{ W (m}^2 \text{ c/s)}^{-1}$

Numbers to the Right of each Column give the Delay time ( $\Delta t$ ) in Minutes.



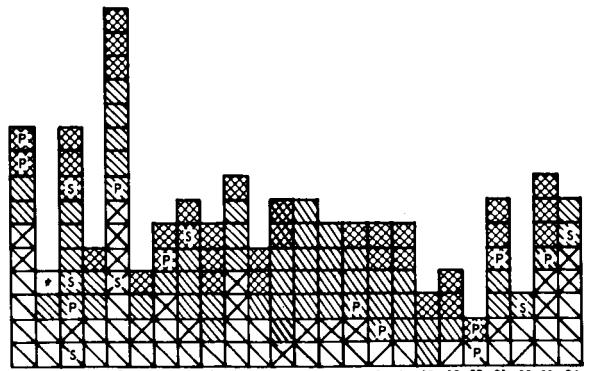


FIGURE 4a TIME DISTRIBUTION IMP  $\leq$  2: FLARES WITH RF FLUX  $\geq 500 \times 10^{-22} \text{ W (m}^2 \text{ c/s})^{-1}$   
Letter P Indicates an Associated Important PCA  
Letter S Indicates an Associated Small or Questionable PCA  
\* Importance 3 not reported in the IAU Bulletin



FIGURE 4.b. TIME DISTRIBUTION RF FLUX  $> 500 \times 10^{-22} \text{ W (m}^2 \text{ c/s})^{-1}$   
NO FLARE PATROL OR NO FLARE REPORTED

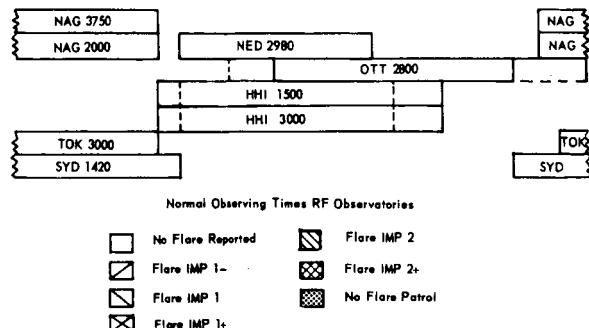
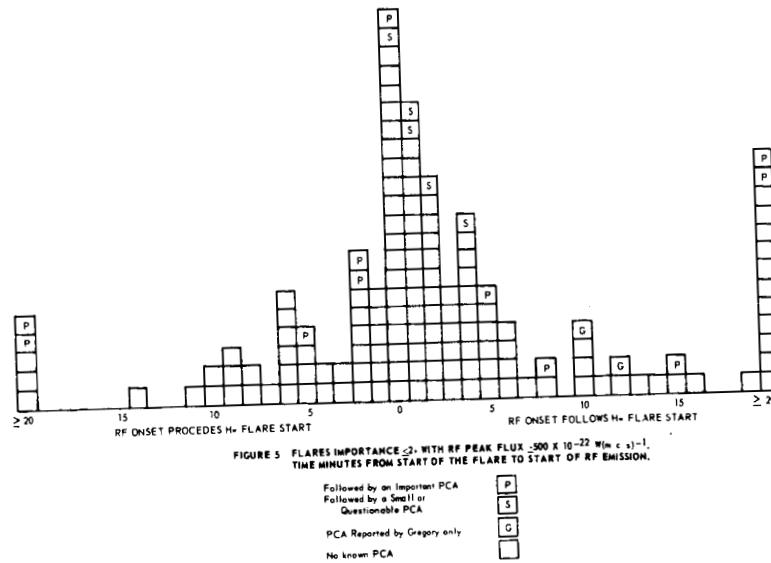


FIGURE 4 TIME DISTRIBUTION RF BURSTS  $\geq 500 \times 10^{-22} \text{ W (m}^2 \text{ c/s})^{-1}$   
WITH ASSOCIATED FLARE IMPORTANCE



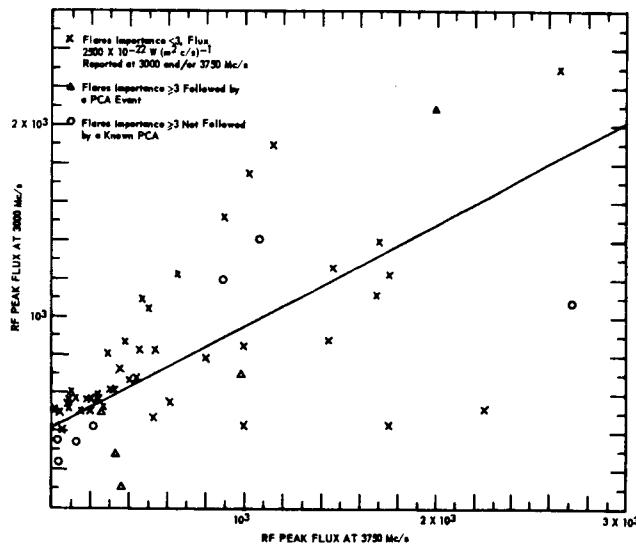


FIGURE 6A - COMPARISON OF RF BURST EMISSIONS AT 3000 Mc/s REPORTED BY TOKYO, AND 3750 Mc/s REPORTED BY NAGOYA

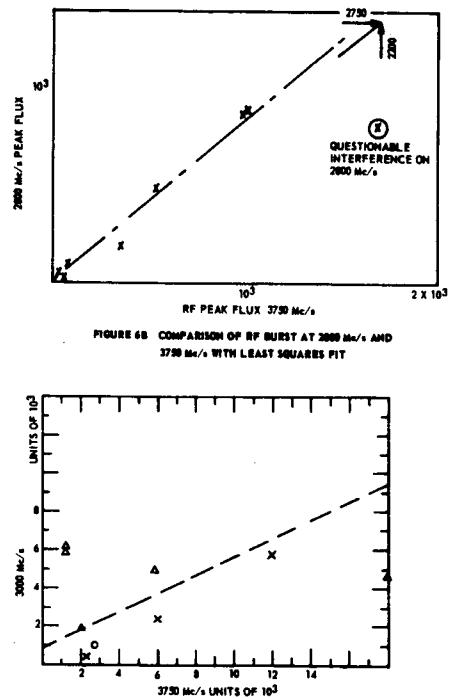
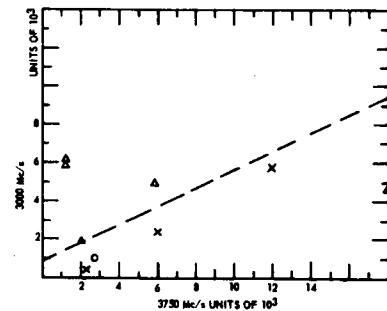


FIGURE 6B - COMPARISON OF RF BURST AT 2800 Mc/s AND 3750 Mc/s WITH LEAST SQUARES FIT



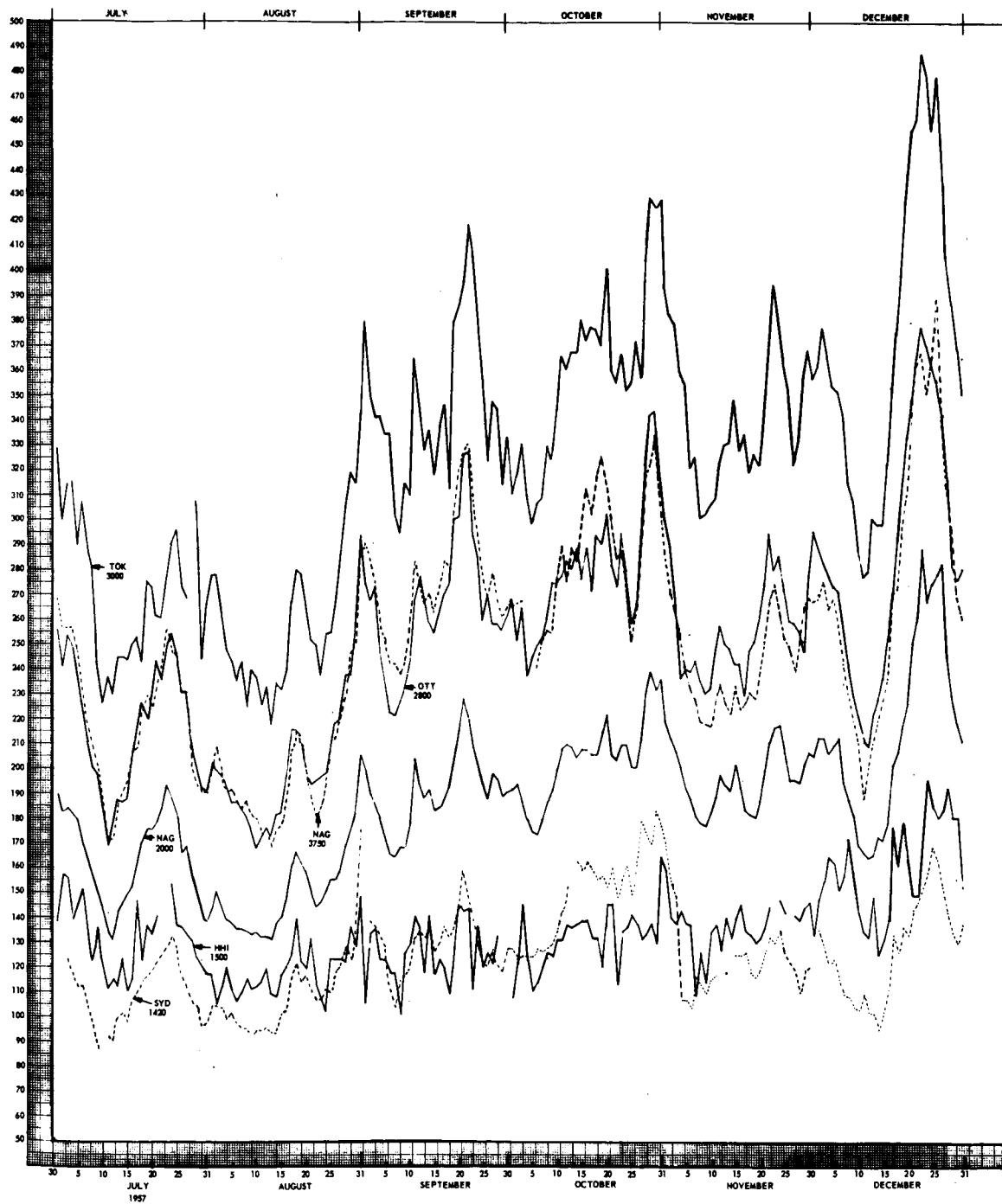
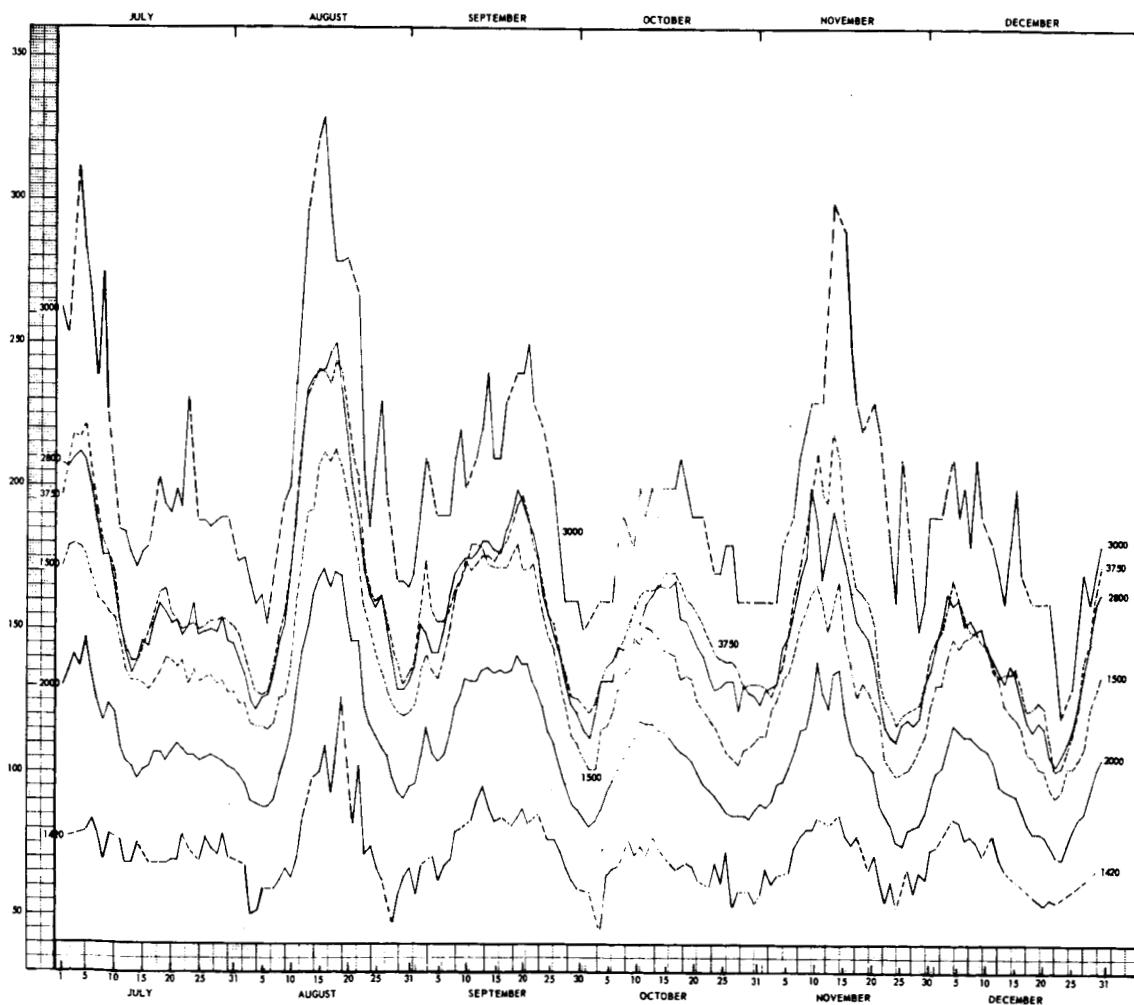


FIGURE 7 COMPARISON OF THE DAILY MEAN RF FLUX AT THE SIX PRINCIPAL FREQUENCIES FOR THE FIRST SIX MONTHS OF THE IGY, INCLUDING THE PERIOD OF SOLAR MAXIMUM



**FIGURE 8 COMPARISON OF THE DAILY MEAN RF FLUX AT THE SIX  
PRINCIPAL FREQUENCIES DURING THE DECLINE OF THE SOLAR CYCLE  
JULY THROUGH DECEMBER 1960**